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# Research of the periodic solutions of the Hénon-Heiles nonintegrable Hamiltonian system by the Lindstedt-Poincaré method 

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#### Abstract

In this paper, we seek the periodic solutions of the Hénon-Heiles nonintegrable Hamiltonian system. We apply the Lindstedt-Poincaré method, in order, first to enumerate the main periodic families in the neighbourhood of the origin, then to determine the series corresponding to these families and to their periods. All the series will be computed to $\mathrm{O}\left(A^{21}\right)$ by means of the computer algebra system 'Mathematica', where $A$ is the zeroth-order amplitude. We also prove that the period of the rectilinear periodic family is exactly equal to a Gauss hypergeometric series. Moreover, we show that the celestial technique of the 'elimination of secular terms' is rigorously equivalent to the 'Fredholm alternative'. We further test the validity of the periodic families using numerical integration. Finally, we compare our results with those of the Churchill-Pecelli-Rod 'geometrical' method.


## 1. Introduction

In this work, we will apply the Lindstedt-Poincaré (LP) method to look for the periodic solutions, in the neighbourhood of the equilibrium point (here the origin), of the HénonHeiles nonintegrable Hamiltonian system whose Hamiltonian is given by

$$
\begin{equation*}
H=\frac{1}{2}\left(\dot{x}^{2}+\dot{y}^{2}+x^{2}+y^{2}\right)+\epsilon\left(x y^{2}-\frac{1}{3} x^{3}\right) \tag{1.1}
\end{equation*}
$$

where $\epsilon$ is a real parameter and $(x, y)$ are the generalized coordinates with $\dot{x}=\frac{\mathrm{d} x}{\mathrm{~d} t}, \dot{y}=\frac{\mathrm{d} y}{\mathrm{~d} t}$.
Using the technique of the stretching variables $\tilde{x}=\epsilon x, \tilde{y}=\epsilon y, \tilde{H}=\epsilon^{2} H$, we can set $\epsilon=1$. The motion of this system is governed by the following differential equations:

$$
\begin{align*}
& \ddot{x}+x=\epsilon\left(x^{2}-y^{2}\right)  \tag{1.2}\\
& \ddot{y}+y=-2 \epsilon x y .
\end{align*}
$$

In view of the theorem of Weinstein-Moser [1], we can assure the existence for this system, for sufficiently small energies, at least two periodic orbits.

Recently, this system aroused the increasing interest of many astronomers, physicists and mathematicians [2-4]. Initially, it was introduced and studied numerically by the two astronomers Hénon and Heiles [5]. Later, it was the subject of numerous numerical and geometrical researches. One of the major problems met with the application of numerical methods to Hamiltonian systems is the accumulation of round-off errors. These last methods also present the inconvenience that they do not offer any information about the number of the periodic families nor about their periods.

[^0]In our work, we shall be interested in the Hénon-Heiles system from a perturbative point of view, using the LP method. We will see that the importance of this method lies in the fact that it permits on the one hand the enumeration of the main periodic families and on the other hand the determination of these families, as well as their periods in the form of perturbative series.

This method is well known in the literature of one-degree-of-freedom systems [6, 7]. It remains very little used in the case of two-degrees-of-freedom systems, in spite of the existence of powerful computer algebra systems. The power of these systems is due to the fact that they handle not only symbolic computations but also numerical computations with any precision.

In section 2, we will describe the LP method in the case of the Hénon-Heiles system, thus deducing a recurrent functional algorithm. We will also prove that the celestial technique of 'elimination of secular terms', on which the LP method is based, is rigorously equivalent to the 'alternative of Fredholm'.

In section 3, we will apply the LP method in order to look for the periodic families of the Hénon-Heiles system in the neighbourhood of the origin. We will then show, exploiting the first steps of the LP algorithm, that the system has eight main periodic families. Actually, taking into account the symmetry and the invariance by rotations of angles $\pm \frac{2 \pi}{3}$ of the potential, we can distinguish only three main periodic families: the 'rectilinear- $\mathcal{R}$ ', the 'curvilinear- $\mathcal{V}$ ' and the 'circular- $\mathcal{C}$ '. By means of the computer algebra system 'Mathematica', we will compute the series corresponding to these periodic families and to their periods to $\mathrm{O}\left(A^{21}\right)$, where $A$ is the zeroth-order amplitude. We will moreover test the validity of these series by numerical integration.

In section 4, we will study the periods of the three main periodic families in terms of the energy $E$. We will first give the coefficients of the power series, truncated to $\mathrm{O}\left(E^{10}\right)$, representing the periods of these families. We will then prove, reducing the system to one degree of freedom and applying a method based on the distributions due to the authors of [8,9], that the period of the rectilinear family is equal to a Gauss hypergeometric series. This will constitute a good check of the rectilinear periodic family computed by the LP method. We will finally discuss the convergence of the series representing the periods of the three main periodic families.

In section 5, we will compare our results with those of the 'geometrical' method of Churchill et al [10]. We will particularly notice the perfect agreement concerning the number and the form of the main periodic families. However, we will point out a disagreement about the circular periodic family.

## 2. Method of Lindstedt-Poincaré

### 2.1. Description of the Lindstedt-Poincaré method

The main purpose of this method is to look for periodic solutions of the system (1.2) in the form

$$
\begin{align*}
& x(t)=\sum_{j=0}^{\infty} x_{j}(t) \cdot \epsilon^{j} \\
& y(t)=\sum_{j=0}^{\infty} y_{j}(t) \cdot \epsilon^{j} \tag{2.1}
\end{align*}
$$

where the functions $x_{j}$ and $y_{j}$ are $T$-periodic. Moreover, the method requires that the pulsation $\omega=\frac{2 \pi}{T}$ is in the form

$$
\begin{equation*}
\omega=\omega_{0}+\omega_{1} \cdot \epsilon+\omega_{2} \cdot \epsilon^{2}+\cdots \tag{2.2}
\end{equation*}
$$

where $\omega_{0}=1$ and $\omega_{j} \in \mathbb{R}$.
Let us make the change of variables

$$
\begin{array}{ll}
\theta=\omega t & \\
X(\theta)=x(t) & X_{j}(\theta)=x_{j}(t)  \tag{2.3}\\
Y(\theta)=y(t) & Y_{j}(\theta)=y_{j}(t)
\end{array}
$$

The system (1.2) then becomes

$$
\begin{align*}
& \omega^{2} X^{\prime \prime}+X=\epsilon\left(X^{2}-Y^{2}\right) \\
& \omega^{2} Y^{\prime \prime}+Y=-2 \epsilon X Y \tag{2.4}
\end{align*}
$$

The problem is now reduced to looking for the $2 \pi$-periodic solutions $(X, Y)$ in the form

$$
\begin{align*}
& X(\theta)=\sum_{j=0}^{\infty} X_{j}(\theta) \cdot \epsilon^{j}  \tag{2.5}\\
& Y(\theta)=\sum_{j=0}^{\infty} Y_{j}(\theta) \cdot \epsilon^{j}
\end{align*}
$$

where the functions $X_{j}$ and $Y_{j}$ are $2 \pi$-periodic.
Setting

$$
\begin{align*}
& F(\theta)=\epsilon\left(X^{2}(\theta)-Y^{2}(\theta)\right)  \tag{2.6}\\
& G(\theta)=-2 \epsilon X(\theta) Y(\theta)
\end{align*}
$$

we obtain

$$
\begin{align*}
& F(\theta)=\sum_{j=0}^{\infty} F_{j}(\theta) \cdot \epsilon^{j}  \tag{2.7}\\
& G(\theta)=\sum_{j=0}^{\infty} G_{j}(\theta) \cdot \epsilon^{j}
\end{align*}
$$

with

$$
\begin{align*}
F_{0} & =G_{0}=0 \\
F_{j} & =\sum_{k=0}^{j-1}\left(X_{k} \cdot X_{j-k-1}-Y_{k} \cdot Y_{j-k-1}\right) \quad j \geqslant 1  \tag{2.8}\\
G_{j} & =-2 \sum_{k=0}^{j-1} X_{k} \cdot Y_{j-k-1} \quad j \geqslant 1 .
\end{align*}
$$

The series $\omega^{2}$ can be written in the form

$$
\begin{align*}
\omega^{2} & =\sum_{j=0}^{\infty} Q_{j} \cdot \epsilon^{j}  \tag{2.9}\\
Q_{j} & =\sum_{k=0}^{j} \omega_{k} \cdot \omega_{j-k}
\end{align*}
$$

Equating the coefficients of $\epsilon^{j}$ on both sides of the system (2.4) we obtain the algorithm

$$
\begin{equation*}
\sum_{k=0}^{j} Q_{j-k}\binom{X_{k}^{\prime \prime}}{Y_{k}^{\prime \prime}}+\binom{X_{j}}{Y_{j}}=\binom{F_{j}}{G_{j}} \quad j \in \mathbb{N} \tag{2.10}
\end{equation*}
$$

which can be written in the form
$\binom{X_{j}^{\prime \prime}+X_{j}}{Y_{j}^{\prime \prime}+Y_{j}}=\binom{F_{j}}{G_{j}}-\sum_{k=1}^{j-1} Q_{j-k}\binom{X_{k}^{\prime \prime}}{Y_{k}^{\prime \prime}}-Q_{j}\binom{X_{0}^{\prime \prime}}{Y_{0}^{\prime \prime}} \quad j \in \mathbb{N}$.
This sequence of second-order linear differential systems of unknowns ( $X_{j}, Y_{j}$ ), constitutes a recurrent functional algorithm permitting in each step $j \in \mathbb{N}$ to simultaneously determine the constant $Q_{j}$ and the solution $\left(X_{j}, Y_{j}\right)$.

Determination of $Q_{j}$ and $\left(X_{j}, Y_{j}\right)$. We are now going to show how to determine recurrently $Q_{j}$ and $\left(X_{j}, Y_{j}\right)$.

Step $j=0$. The system of this step is

$$
\begin{align*}
& X_{0}^{\prime \prime}+X_{0}=0 \\
& Y_{0}^{\prime \prime}+Y_{0}=0 \tag{2.12}
\end{align*}
$$

whose solution is given by

$$
\begin{align*}
& X_{0}=A \cos \theta+B \sin \theta \\
& Y_{0}=C \cos \theta+D \sin \theta \tag{2.13}
\end{align*}
$$

where $A, B, C$ and $D$ are arbitrary constants of integration.
We also have here

$$
\begin{equation*}
Q_{0}=1 \tag{2.14}
\end{equation*}
$$

Step $j=1$. We have here the system

$$
\begin{equation*}
\binom{X_{1}^{\prime \prime}+X_{1}}{Y_{1}^{\prime \prime}+Y_{1}}=\binom{F_{1}}{G_{1}}-Q_{1}\binom{X_{0}^{\prime \prime}}{Y_{0}^{\prime \prime}} \tag{2.15}
\end{equation*}
$$

where

$$
\begin{equation*}
\binom{F_{1}}{G_{1}}=\binom{X_{0}^{2}-Y_{0}^{2}}{-2 X_{0} Y_{0}} . \tag{2.16}
\end{equation*}
$$

After some calculations we obtain

$$
\binom{F_{1}}{G_{1}}(\theta)=\binom{a_{10}}{c_{10}}+\sum_{k=1}^{2}\left(\begin{array}{ll}
a_{1 k} & b_{1 k}  \tag{2.17}\\
c_{1 k} & d_{1 k}
\end{array}\right)\binom{\cos k \theta}{\sin k \theta}
$$

with

$$
\begin{align*}
& a_{10}=\frac{A^{2}+B^{2}-C^{2}-D^{2}}{2} \quad c_{10}=-A C-B D \\
& a_{11}=0 \quad c_{11}=0 \\
& a_{12}=\frac{A^{2}+D^{2}-B^{2}-C^{2}}{2} \quad c_{12}=B D-A C  \tag{2.18}\\
& b_{11}=0 \quad d_{11}=0 \\
& b_{12}=A B-C D \quad d_{12}=-A D-B C .
\end{align*}
$$

The system (2.15) then becomes

$$
\begin{align*}
\binom{X_{1}^{\prime \prime}+X_{1}}{Y_{1}^{\prime \prime}+Y_{1}} & =\binom{a_{10}}{c_{10}}+\left(\begin{array}{ll}
a_{11}+A \cdot Q_{1} & b_{11}+B \cdot Q_{1} \\
c_{11}+C \cdot Q_{1} & d_{11}+D \cdot Q_{1}
\end{array}\right)\binom{\cos \theta}{\sin \theta} \\
& +\left(\begin{array}{ll}
a_{12} & b_{12} \\
c_{12} & d_{12}
\end{array}\right)\binom{\cos 2 \theta}{\sin 2 \theta} \tag{2.19}
\end{align*}
$$

Determination of $Q_{1}$. In order to avoid the secular terms $(\theta \cos \theta$ and $\theta \sin \theta)$ in $X_{1}$ and $Y_{1}$, the coefficients of $\cos \theta$ and $\sin \theta$ have to vanish in the system (2.19).

So we obtain four algebraic equations

$$
\begin{array}{ll}
a_{11}+A \cdot Q_{1}=0 & c_{11}+C \cdot Q_{1}=0 \\
b_{11}+B \cdot Q_{1}=0 & d_{11}+D \cdot Q_{1}=0 \tag{2.20}
\end{array}
$$

The constant $Q_{1}$ exists then if and only if the six determinants

$$
\begin{array}{ll}
D_{1}(A, B)=B a_{11}-A b_{11} & D_{1}(B, C)=C b_{11}-B c_{11} \\
D_{1}(A, C)=C a_{11}-A c_{11} & D_{1}(B, D)=D b_{11}-B d_{11}  \tag{2.21}\\
D_{1}(A, D)=D a_{11}-A d_{11} & D_{1}(C, D)-D c_{11}-C d_{11}
\end{array}
$$

are null. In our case they vanish since we have

$$
a_{11}=b_{11}=c_{11}=d_{11}=0
$$

We can then deduce the constant $Q_{1}$ from one of the equations (2.20)

$$
\begin{equation*}
Q_{1}=-\frac{a_{11}}{A}=0 \quad(\text { if } A \neq 0) \tag{2.22}
\end{equation*}
$$

Determination of $X_{1}$ and $Y_{1}$. After eliminating the terms $\cos \theta$ and $\sin \theta$, the resolution of the system (2.19) gives
$\binom{X_{1}}{Y_{1}}=\binom{A_{1} B_{1}}{C_{1} D_{1}}\binom{\cos \theta}{\sin \theta}+\binom{a_{10}}{c_{10}}+\sum_{k=2}^{2}\left(\frac{1}{1-k^{2}}\right)\left(\begin{array}{ll}a_{1 k} & b_{1 k} \\ c_{1 k} & d_{1 k}\end{array}\right)\binom{\cos k \theta}{\sin k \theta}$
where $A_{1}, B_{1}, C_{1}$ and $D_{1}$ are arbitrary constants of integration. Since we have not yet obtained a relation between $A, B, C$ and $D$, we can set $A_{1}=B_{1}=C_{1}=D_{1}=0$.

Step $j$. The system of this step is

$$
\begin{equation*}
\binom{X_{j}^{\prime \prime}+X_{j}}{Y_{j}^{\prime \prime}+Y_{j}}=\binom{F_{j}}{G_{j}}-\sum_{k=1}^{j-1} Q_{j-k}\binom{X_{k}^{\prime \prime}}{Y_{k}^{\prime \prime}}-Q_{j}\binom{X_{0}^{\prime \prime}}{Y_{0}^{\prime \prime}} \tag{2.24}
\end{equation*}
$$

Now suppose, by hypothesis of induction, we have determined $Q_{i}$ and $\left(X_{i}, Y_{i}\right)$ for all $0 \leqslant i \leqslant j-1$ with

$$
\binom{X_{i}}{Y_{i}}=\left(\begin{array}{cc}
A_{i} & B_{i}  \tag{2.25}\\
C_{i} & D_{i}
\end{array}\right)\binom{\cos \theta}{\sin \theta}+\binom{a_{i 0}}{c_{i 0}}+\sum_{k=2}^{i+1}\left(\frac{1}{1-k^{2}}\right)\left(\begin{array}{cc}
a_{i k} & b_{i k} \\
c_{i k} & d_{i k}
\end{array}\right)\binom{\cos k \theta}{\sin k \theta}
$$

where $A_{i}, B_{i}, C_{i}, D_{i}$, are arbitrary constants of integration.
We can then prove easily that at this step $j$ we have

$$
\binom{F_{j}}{G_{j}}-\sum_{k=1}^{j-1} Q_{j-k}\binom{X_{k}^{\prime \prime}}{Y_{k}^{\prime \prime}}=\binom{a_{j 0}}{c_{j 0}}+\sum_{k=2}^{j+1}\left(\begin{array}{cc}
a_{j k} & b_{j k}  \tag{2.26}\\
c_{j k} & d_{j k}
\end{array}\right)\binom{\cos k \theta}{\sin k \theta}
$$

where $a_{j k}, b_{j k}, c_{j k}$ and $d_{j k}$ depend on the constants $A, B, C$ and $D$.

The system (2.24) then becomes

$$
\begin{align*}
\binom{X_{j}^{\prime \prime}+X_{j}}{Y_{j}^{\prime \prime}+Y_{j}}= & \binom{a_{j 0}}{c_{j 0}}+\left(\begin{array}{cc}
a_{j 1}+A \cdot Q_{j} & b_{j 1}+B \cdot Q_{j} \\
c_{j 1}+C \cdot Q_{j} & d_{j 1}+D \cdot Q_{j}
\end{array}\right)\binom{\cos \theta}{\sin \theta} \\
& +\sum_{k=2}^{j+1}\left(\begin{array}{cc}
a_{j k} b_{j k} & \\
c_{j k} & d_{j k}
\end{array}\right)\binom{\cos k \theta}{\sin k \theta} \tag{2.27}
\end{align*}
$$

Determination of $Q_{j}$. Using the same technique of 'elimination of secular terms' applied in the step $j=1$, we obtain four algebraic equations

$$
\begin{align*}
a_{j 1}+A \cdot Q_{j}=0 & c_{j 1}+C \cdot Q_{j}=0  \tag{2.28}\\
b_{j 1}+B \cdot Q_{j}=0 & d_{j 1}+D \cdot Q_{j}=0
\end{align*}
$$

The constant $Q_{j}$ exists then if and only if the six determinants

$$
\begin{array}{lc}
D_{j}(A, B)=B a_{j 1}-A b_{j 1} & D_{j}(B, C)=C b_{j 1}-B c_{j 1} \\
D_{j}(A, C)=C a_{j 1}-A c_{j 1} & D_{j}(B, D)=D b_{j 1}-B d_{j 1}  \tag{2.29}\\
D_{j}(A, D)=D a_{j 1}-A d_{j 1} & D_{j}(C, D)=D c_{j 1}-C d_{j 1}
\end{array}
$$

are null.
Setting these determinants equal to zero permits us to eventually find a relation between the constants of integration introduced in the preceding steps. To determine $Q_{j}$ it is sufficient to use one of the equations (2.28)

$$
\begin{equation*}
Q_{j}=-\frac{a_{j 1}}{A}(\text { if } A \neq 0) \tag{2.30}
\end{equation*}
$$

Determination of $X_{j}$ and $Y_{j}$. After eliminating the terms $\cos (\theta)$ and $\sin (\theta)$, the resolution of the system (2.24) gives

$$
\binom{X_{j}}{Y_{j}}=\left(\begin{array}{cc}
A_{j} & B_{j}  \tag{2.31}\\
C_{j} & D_{j}
\end{array}\right)\binom{\cos \theta}{\sin \theta}+\binom{a_{j 0}}{c_{j 0}}+\sum_{k=2}^{j+1}\left(\frac{1}{1-k^{2}}\right)\left(\begin{array}{cc}
a_{j k} & b_{j k} \\
c_{j k} & d_{j k}
\end{array}\right)\binom{\cos k \theta}{\sin k \theta}
$$

where $A_{j}, B_{j}, C_{j}$ and $D_{j}$ are arbitrary constants of integration.

### 2.2. Fredholm alternative and secular terms

We will now show that the well known celestial technique of 'elimination of secular terms', on which the LP method is based, is rigorously equivalent to the Fredholm alternative.

Let us now rewrite the linear differential system (2.24) in the form

$$
\begin{equation*}
\binom{X_{j}^{\prime \prime}}{Y_{j}^{\prime \prime}}+\binom{X_{j}}{Y_{j}}=\binom{f_{j}}{g_{j}} \tag{2.32}
\end{equation*}
$$

where

$$
\begin{gather*}
\binom{f_{j}}{g_{j}}=\binom{a_{j 0}}{c_{j 0}}+\binom{a_{j 1}+A Q_{j}}{c_{j 1}+C Q_{j}} \cos \theta+\binom{b_{j 1}+B Q_{j}}{d_{j 1}+D Q_{j}} \sin \theta \\
+\sum_{k=2}^{j+1}\left(\begin{array}{cc}
a_{j k} & b_{j k} \\
c_{j k} & d_{j k}
\end{array}\right)\binom{\cos k \theta}{\sin k \theta} \tag{2.33}
\end{gather*}
$$

Theorem. (Fredholm alternative, cf Reinhard [11, p 394]). Let us consider the differential equation in $\mathbb{R}^{d}$

$$
\text { (I) } \quad \frac{\mathrm{d} x}{\mathrm{~d} t}=A(t) x+b(t)
$$

where the functions $b$ and $A$, defined in $\mathbb{R}$ and with values respectively in $\mathbb{R}^{d}$ and $\mathcal{L}\left(\mathbb{R}^{d}\right)$, are continuous and $T$-periodic. We assume that the following adjoint equation of the homogeneous equation

$$
\begin{equation*}
\frac{\mathrm{d} y}{\mathrm{~d} t}=-{ }^{t} A(t) \cdot y \tag{II}
\end{equation*}
$$

admits $p$ independent $T$-periodic solutions $y_{1}, y_{2}, \ldots, y_{p}$. Then there exists $p$ independent $T$-periodic solutions for the equation (I) if and only if for all $k \in \mathbb{N}$ such that $k \leqslant p$ we have

$$
\begin{equation*}
\int_{0}^{T}\left\langle y_{k}(t), b(t)\right\rangle \mathrm{d} t=0 \tag{III}
\end{equation*}
$$

where $\langle\cdot, \cdot\rangle$ denotes the scalar product in $\mathbb{R}^{d}$.
Let us apply this theorem to the system (2.32). If we set

$$
\begin{align*}
& u_{1}=X_{j} \\
& u_{2}=X_{j}^{\prime}  \tag{2.34}\\
& u_{3}=Y_{j} \\
& u_{4}=Y_{j}^{\prime}
\end{align*}
$$

then the system (2.32) can be written in the normal form

$$
\left(\begin{array}{l}
u_{1}^{\prime}  \tag{2.35}\\
u_{2}^{\prime} \\
u_{3}^{\prime} \\
u_{4}^{\prime}
\end{array}\right)=\left(\begin{array}{cccc}
0 & 1 & 0 & 0 \\
-1 & 0 & 0 & 0 \\
0 & 0 & 0 & 1 \\
0 & 0 & -1 & 0
\end{array}\right)\left(\begin{array}{l}
u_{1} \\
u_{2} \\
u_{3} \\
u_{4}
\end{array}\right)+\left(\begin{array}{c}
0 \\
f_{j} \\
0 \\
g_{j}
\end{array}\right)
$$

or equivalently in the condensed form

$$
\begin{equation*}
u^{\prime}=A u+b . \tag{2.36}
\end{equation*}
$$

It is clear that $A$ and $b$ are $2 \pi$-periodic. The adjoint equation of (2.36) is then given by

$$
\begin{equation*}
v^{\prime}=A v \tag{2.37}
\end{equation*}
$$

We can easily prove that this equation admits the four following independent $2 \pi$-periodic solutions
$v_{1}=\left(\begin{array}{c}\cos \theta \\ -\sin \theta \\ 0 \\ 0\end{array}\right) \quad v_{2}=\left(\begin{array}{c}\sin \theta \\ \cos \theta \\ 0 \\ 0\end{array}\right) \quad v_{3}=\left(\begin{array}{c}0 \\ 0 \\ \cos \theta \\ -\sin \theta\end{array}\right) \quad v_{4}=\left(\begin{array}{c}0 \\ 0 \\ \sin \theta \\ \cos \theta\end{array}\right)$
which are the rows of the resolvant $\mathrm{e}^{A \theta}$.
In view of the preceding theorem, the system (2.36) admits $2 \pi$-periodic solutions if and only if for all $k \in \mathbb{N}$ such that $1 \leqslant k \leqslant 4$ we have

$$
\begin{equation*}
\int_{0}^{2 \pi}\left\langle v_{k}(\theta), b(\theta)\right\rangle \mathrm{d} \theta=0 \tag{2.39}
\end{equation*}
$$

i.e.

$$
\begin{align*}
& \int_{0}^{2 \pi}\binom{f_{j}(\theta)}{g_{j}(\theta)} \cos \theta \mathrm{d} \theta=\binom{0}{0}  \tag{2.40}\\
& \int_{0}^{2 \pi}\binom{f_{j}(\theta)}{g_{j}(\theta)} \sin \theta \mathrm{d} \theta=\binom{0}{0}
\end{align*} .
$$

The first equality is equivalent to

$$
\begin{gather*}
\binom{a_{j 0}}{c_{j 0}} \int_{0}^{2 \pi} \cos \theta \mathrm{~d} \theta+\binom{a_{j 1}+A Q_{j}}{c_{j 1}+C Q_{j}} \int_{0}^{2 \pi} \cos ^{2} \theta \mathrm{~d} \theta+\binom{b_{j 1}+B Q_{j}}{d_{j 1}+D Q_{j}} \int_{0}^{2 \pi} \sin \theta \cdot \cos \theta \mathrm{~d} \theta \\
\quad+\sum_{k=2}^{j+1}\left(\begin{array}{cc}
a_{j k} & b_{j k} \\
c_{j k} & d_{j k}
\end{array}\right)\binom{\int_{0}^{2 \pi} \cos k \theta \cos \theta \mathrm{~d} \theta}{\int_{0}^{2 \pi} \sin k \theta \cos \theta \mathrm{~d} \theta} \\
=\binom{0}{0} \tag{2.41}
\end{gather*}
$$

It is evident that all the integrals in the relation (2.41) are null except $\int_{0}^{2 \pi} \cos ^{2} \theta \mathrm{~d} \theta=\pi$. The relation (2.41) is therefore equivalent to two algebraic equations

$$
\begin{align*}
& a_{j 1}+A Q_{j}=0  \tag{2.42}\\
& c_{j 1}+C Q_{j}=0
\end{align*}
$$

Proceeding in the same manner for the second equality of (2.40), we obtain two other algebraic equations

$$
\begin{align*}
b_{j 1}+B Q_{j} & =0 \\
d_{j 1}+D Q_{j} & =0 \tag{2.43}
\end{align*}
$$

We thus recover the four algebraic equations (2.28) obtained using the technique of 'elimination of secular terms'.

## 3. Enumeration and computation of the periodic families of the Hénon-Heiles system

### 3.1. Enumeration of the main periodic families of the Hénon-Heiles system

We will see here that the enumeration of the main periodic families is accomplished at the fourth step of the algorithm (2.11).

Step $j=0$. The solution of this step is

$$
\begin{align*}
& X_{0}=A \cos \theta+B \sin \theta \\
& Y_{0}=C \cos \theta+D \sin \theta  \tag{3.1}\\
& Q_{0}=1
\end{align*}
$$

Step $j=1$. The solution here is

$$
\begin{align*}
& X_{1}=a_{10}-\frac{1}{3}\left(a_{12} \cos 2 \theta+b_{12} \sin 2 \theta\right. \\
& Y_{1}=c_{10}-\frac{1}{3}\left(c_{12} \cos 2 \theta+d_{12} \sin 2 \theta\right)  \tag{3.2}\\
& Q_{1}=0
\end{align*}
$$

where the coefficients $a_{1 j}, b_{1 j}, c_{1 j}$ and $d_{1 j}$ are given in (2.18).

Step $j=2$. To compute the six determinants $D_{2}$ of this step we must calculate the coefficients $a_{21}, b_{21}, c_{21}$ and $d_{21}$

$$
\begin{align*}
& a_{21}=\frac{5}{6}\left(A^{3}+A B^{2}+A C^{2}\right)+\frac{7}{3} B C D-\frac{3}{2} A D^{2} \\
& b_{21}=\frac{5}{6}\left(B^{3}+B A^{2}+B D^{2}\right)+\frac{7}{3} A C D-\frac{3}{2} B C^{2}  \tag{3.3}\\
& c_{21}=\frac{5}{6}\left(C^{3}+C D^{2}+C A^{2}\right)+\frac{7}{3} A B D-\frac{3}{2} C B^{2} \\
& d_{21}=\frac{5}{6}\left(D^{3}+D C^{2}+D B^{2}\right)+\frac{7}{3} A B C-\frac{3}{2} D A^{2}
\end{align*}
$$

We then deduce the six determinants

$$
\begin{array}{lc}
D_{2}(A, B)=\frac{7}{3}(A C+B D)(B C-A D) & \left.D_{2}(B, C)=\frac{7}{3}\left(B^{2}-C^{2}\right) B C-A D\right) \\
D_{2}(A, C)=\frac{7}{3}(A B+C D)(B C-A D) & D_{2}(B, D)=\frac{7}{3}(A B+C D)(A D-B C)  \tag{3.4}\\
D_{2}(A, D)=\frac{7}{3}\left(A^{2}-D^{2}\right)(A D-B C) & D_{2}(C, D)=\frac{7}{3}(A C+B D)(A D-B C)
\end{array}
$$

which have to vanish. We can distinguish two cases.

$$
\begin{array}{lcrrr}
\text { First case } & (A, B)=(0,0) & \text { or } & (C, D)=(0,0) \\
\text { Second case } & (A, B) \neq(0,0) & \text { and } & (C, D) \neq(0,0) .
\end{array}
$$

The zeroth-order solution given in (3.1) may be written as

$$
\begin{align*}
& X_{0}=\alpha_{1} \cos \left(\theta+\theta_{1}\right) \\
& Y_{0}=\alpha_{2} \cos \left(\theta+\theta_{2}\right) \tag{3.5}
\end{align*}
$$

First case. $\quad(A, B)=(0,0)$ or $(C, D)=(0,0)$
(i) $(A, B)=(0,0)$. Thus

$$
\begin{aligned}
& X_{0}=0 \\
& Y_{0}=\alpha_{2} \cos \left(\theta+\theta_{2}\right)
\end{aligned}
$$

As the system is autonomous, we can make the change of variable $\theta \rightarrow \theta-\theta_{2}$. We then obtain

$$
\begin{align*}
& X_{0}=0 \\
& Y_{0}=C \cos \theta \tag{3.6}
\end{align*}
$$

In this case all six determinants (3.4) are null without any condition.
(ii) $(C, D)=(0,0$. Thus

$$
\begin{aligned}
& X_{0}=\alpha_{1} \cos \left(\theta+\theta_{1}\right) \\
& Y_{0}=0
\end{aligned}
$$

which may be written, after the change of variable $\theta \rightarrow \theta-\theta_{1}$, in the form

$$
\begin{align*}
& X_{0}=A \cos \theta \\
& Y_{0}=0 \tag{3.7}
\end{align*}
$$

In this case we also remark that the six determinants (3.4) are all null without any condition.

Second case. $\quad(A, B) \neq(0,0)$ and $(C, D) \neq(0,0)$. In this case we have

$$
\begin{array}{ll}
\alpha_{1}=\sqrt{A^{2}+B^{2}} & \alpha_{2}=\sqrt{C^{2}+D^{2}} \\
\sin \theta_{1}=-\frac{B}{\alpha_{1}} & \sin \theta_{2}=-\frac{D}{\alpha_{2}}  \tag{3.8}\\
\cos \theta_{1}=\frac{A}{\alpha_{1}} & \cos \theta_{2}=\frac{C}{\alpha_{2}}
\end{array}
$$

If we equate the determinant $D_{2}(A, B)$ given in (3.4) to zero, we obtain

$$
\begin{equation*}
B C-A D=0 \quad \text { or } \quad A C+B D=0 \tag{3.9}
\end{equation*}
$$

which is equivalent to

$$
\begin{array}{lll}
\sin \left(\theta_{2}-\theta_{1}\right)=0 & \text { or } & \cos \left(\theta_{2}-\theta_{1}\right)=0 \\
\Leftrightarrow \theta_{2}-\theta_{1}=k \pi & \text { or } & \theta_{2}-\theta_{1}=\frac{\pi}{2}+k \pi \tag{3.10}
\end{array}
$$

Hence $\left(\theta_{2}-\theta_{1}\right) \in\left\{0, \frac{\pi}{2}, \pi, \frac{3 \pi}{2}\right\}$.
As the system is autonomous, we can make the change of variable $\theta \rightarrow \theta-\theta_{1}$. We thus have

$$
\begin{align*}
& X_{0}=\alpha_{1} \cos \theta  \tag{3.11}\\
& Y_{0}=\alpha_{2} \cos \left(\theta+\theta_{2}-\theta_{1}\right)
\end{align*}
$$

Since we have $\left(\theta_{2}-\theta_{1}\right) \in\left\{0, \frac{\pi}{2}, \pi, \frac{3 \pi}{2}\right\}$, we can distinguish two cases

$$
\begin{array}{lll}
X_{0}=A \cos \theta & \text { and } & X_{0}=A \cos \theta \\
Y_{0}=C \cos \theta & \text { and } & Y_{0}=D \sin \theta \tag{3.12}
\end{array}
$$

It is clear that the previous cases (3.6) and (3.7) are included in (3.12) by setting $A=0$ or $C=0$.

Conclusion. The autonomy of the system and the resolution of the equation $D_{2}(A, B)=0$ imply that (3.12) are the only possible forms of ( $X_{0}, Y_{0}$ ).

We have now to equate to zero, for each case of (3.12), the five others determinants of (3.4).

First case $\left\{\begin{array}{l}X_{0}=A \cos \theta \\ Y_{0}=C \cos \theta\end{array} \quad\right.$ (here we have $B=D=0$ ).
In this case all the determinants of (3.4) are null without any condition on $A$ and $C$. We then go to the step $j=3$.

Step $j=3$. After some calculations we obtain

$$
\begin{equation*}
a_{31}=b_{31}=c_{31}=d_{31}=0 \tag{3.13}
\end{equation*}
$$

It follows that the six determinants of this step are null without any condition on $A$ and $C$. We then go to the step $j=4$.

Step $j=4$. We obtain after some calculations

$$
\begin{align*}
& a_{41}=2 A\left(a_{30}-\frac{1}{6} a_{32}\right)-2 C\left(c_{30}-\frac{1}{6} c_{32}\right)+\frac{1}{24}\left(a_{12} a_{23}-c_{12} c_{23}\right) \\
& c_{41}=-2 A\left(c_{30}-\frac{1}{6} c_{32}\right)-2 C\left(a_{30}-\frac{1}{6} a_{32}\right)-\frac{1}{24}\left(a_{12} c_{23}+a_{23} c_{12}\right)  \tag{3.14}\\
& b_{41}=d_{41}=0
\end{align*}
$$

with
$a_{30}=\frac{19}{72}\left(A^{4}+C^{4}-6 A^{2} C^{2}\right) \quad c_{30}=\frac{19}{18} A C\left(A^{2}-C^{2}\right)$
$a_{32}=\frac{1}{144}\left(59 A^{4}-107 C^{4}+144 A^{2} C^{2}\right) \quad c_{32}=-\frac{1}{72} A C\left(131 A^{2}+35 C^{2}\right)$
$a_{34}=\frac{1}{144}\left(5 A^{4}-C^{4}-12 A^{2} C^{2}\right) \quad c_{32}=\frac{1}{72} A C\left(A^{2}-7 C^{2}\right)$
and $a_{12}, c_{12}$ are given in (2.18).
We can remark here that all the determinants $D_{4}$ are null except $D_{4}(A, C)$. The calculation of this determinant leads to

$$
\begin{equation*}
D_{4}(A, C)=\frac{7}{6} A C\left(3 A^{4}-10 A^{2} C^{2}+3 C^{4}\right) \tag{3.16}
\end{equation*}
$$

which vanish if

$$
\begin{equation*}
A=0 \quad \text { or } \quad C=0 \quad \text { or } \quad 3 A^{4}-10 A^{2} C^{2}+3 C^{4}=0 \tag{3.17}
\end{equation*}
$$

or equivalently if
$A=0 \quad$ or $\quad C=0 \quad$ or $\quad C= \pm \sqrt{3} A \quad$ or $\quad C= \pm \frac{A}{\sqrt{3}}$.
We conclude that the only possible forms of this first case are

$$
\begin{align*}
& X_{0}=0 \quad X_{0}=A \cos \theta \quad X_{0}=A \cos \theta \quad X_{0}=A \cos \theta \\
& Y_{0}=C \cos \theta \quad Y_{0}=0 \quad Y_{0}= \pm \sqrt{3} A \cos \theta \quad Y_{0}= \pm \frac{A}{\sqrt{3}} \cos \theta  \tag{3.19}\\
& \text { second case }\left\{\begin{array}{l}
X_{0}=A \cos \theta \\
Y_{0}=D \sin \theta
\end{array} \quad \text { (we have here } B=C=0\right. \text { ). }
\end{align*}
$$

We remark here that the six determinants $D_{2}$ of the step 2 are null except the determinant

$$
\begin{equation*}
D_{2}(A, D)=\frac{7}{3} A D\left(A^{2}-D^{2}\right) \tag{3.20}
\end{equation*}
$$

which is null if

$$
\begin{equation*}
A=0 \quad \text { or } \quad D=0 \quad \text { or } \quad D= \pm A \tag{3.21}
\end{equation*}
$$

We deduce that the only possible forms of this second case are

$$
\begin{array}{lrr}
X_{0}=0 & X_{0}=A \cos \theta & X_{0}=A \cos \theta \\
Y_{0}=D \sin \theta & Y_{0}=0 & Y_{0}= \pm A \sin \theta \tag{3.22}
\end{array}
$$

General conclusion. The algorithm can be started with one of the following eight possible cases
$X_{0}=0 \quad X_{0}=A \cos \theta \quad X_{0}=A \cos \theta \quad X_{0}=A \cos \theta \quad X_{0}=A \cos \theta$
$Y_{0}=C \cos \theta \quad Y_{0}=0 \quad Y_{0}= \pm \sqrt{3} A \cos \theta \quad Y_{0}= \pm \frac{A}{\sqrt{3}} \cos \theta$
$Y_{0}= \pm A \sin \theta$.
Actually, each case corresponds to a periodic family parametrized by the zeroth-order amplitude $A$. We conclude then that the LP method permits the determination of eight main


Figure 1. (a) Three rectilinear periodic orbits computed by the LP method. (b) Three rectilinear periodic orbits computed by numerical integration.
periodic families for the Hénon-Heiles system. Taking into account the symmetry and the invariance by rotations of angles $\pm \frac{2 \pi}{3}$ of the potential, we can in fact distinguish only three main periodic families: the 'rectilinear- $\mathcal{R}$ ', the 'curvilinear- $\mathcal{V}$ ' and the 'circular- $\mathcal{C}$ '.

Below we give the series, truncated to $\mathrm{O}\left(A^{21}\right)$, representing these three families and their periods.

### 3.2. Computation by LP and test by numerical integration of the main periodic families

As mentioned in the introduction, we can set $\epsilon=1$ and we consequently parametrize the families by the parameter $A$ or by the energy $E$ (constant of motion) which is related to $A$ by the relation

$$
\begin{equation*}
E=\sum_{j=1}^{\infty} E_{j} A^{2 j} . \tag{3.24}
\end{equation*}
$$

By means of the computer algebra system 'Mathematica', we obtain with exact calculation the three main periodic families as well as their periods to $\mathrm{O}\left(A^{21}\right)$.

In figures 1-3 we compare, for the energy $E=\frac{1}{8}$, the periodic orbits computed by the LP method, with those obtained by numerical integration.
(1) 'Rectilinear' periodic family $\mathcal{R}$. This family contains in fact three straight periodic families: the 'horizontal' $\mathcal{H}(y=0)$ and two 'oblique' $\mathcal{O}^{ \pm}(y= \pm \sqrt{3} x)$.

Below we give (3.25) the series representing the 'horizontal' family; the two oblique can be deduced from the previous family by rotations of angles $\pm \frac{2 \pi}{3}$.

$$
\begin{aligned}
X(\theta)=A \cos (\theta) & +A^{2}\left(\frac{1}{2}-\frac{\cos (2 \theta)}{6}\right)+A^{3} \frac{\cos (3 \theta)}{48} \\
& +A^{4}\left(\frac{19}{79}-\frac{59 \cos (2 \theta)}{432}-\frac{\cos (4 \theta)}{432}\right)+A^{5}\left(\frac{79 \cos (3 \theta)}{2304}+\frac{5 \cos (5 \theta)}{20736}\right) \\
& +A^{6}\left(\frac{11897}{41472}-\frac{1207 \cos (2 \theta)}{6912}-\frac{89 \cos (4 \theta)}{15552}-\frac{\cos (6 \theta)}{41472}\right)
\end{aligned}
$$



Figure 2. (a) Three curvilinear periodic orbits computed by the LP method. (b) Three curvilinear periodic orbits computed by numerical integration. (Energy $E=\frac{1}{8}$.)


Figure 3. (a) Circular periodic orbit computed by the LP method. (b) Circular periodic orbit computed by numerical integration. (Energy $E=\frac{1}{8}$.)

$$
\begin{align*}
& +A^{7}\left(\frac{19283 \cos (3 \theta)}{331776}+\frac{2375 \cos (5 \theta)}{2985984}+\frac{7 \cos (7 \theta)}{2985984}\right) \\
& +A^{8}\left(\frac{1181413}{2985984}-\frac{1151545 \cos (2 \theta)}{4478976}-\frac{54067 \cos (4 \theta)}{4478976}\right. \\
& \left.-\frac{11 \cos (6 \theta)}{110592}-\frac{\cos (8 \theta)}{4478976}\right) \\
& +A^{9}\left(\frac{537439 \cos (3 \theta)}{5308416}+\frac{215725 \cos (5 \theta)}{107495424}+\frac{4991 \cos (7 \theta)}{429981696}+\frac{\cos (9 \theta)}{47775744}\right) \\
& +A^{10}\left(\frac{132341179}{214990848}-\frac{528929135 \cos (2 \theta)}{1289945088}-\frac{1305943 \cos (4 \theta)}{53747712}\right.  \tag{3.25}\\
& \left.-\frac{83729 \cos (6 \theta)}{286654464}-\frac{13 \cos (8 \theta)}{10077696}\right)+\cdots
\end{align*}
$$

$$
\begin{aligned}
& +A^{21}\left(\frac{1846201729551096763039 \cos (3 \theta)}{425973332494163902464}+\cdots\right. \\
& \left.+\frac{7 \cos (21 \theta)}{1277919997482491707392}\right)+\mathrm{O}\left(A^{21}\right)
\end{aligned}
$$

$Y(\theta)=0$

$$
\begin{aligned}
\omega^{2}=1-\frac{5}{6} A^{2} & -\frac{335}{864} A^{4}-\frac{16195}{41472} A^{6}-\frac{9244585}{17915904} A^{8}-\frac{505885685}{644972544} A^{10} \\
& -\frac{53471731165}{41278242816} A^{12}-\frac{120752481301295}{53496602689536} A^{14} \\
& -\frac{31442429175813815}{7703510787293184} A^{16}-\frac{8411829860794238695}{1109305553370218496} A^{18} \\
& -\frac{1148709686193583415155}{798699998426555731712} A^{20}+\mathrm{O}\left(A^{21}\right)
\end{aligned}
$$

(2) 'Curvilinear' periodic family $\mathcal{V}$. Actually, this family contains three curvilinear families $\mathcal{V}_{1}, \mathcal{V}_{2}$ and $\mathcal{V}_{3}$. Below we give (3.26) the series representing the family $\mathcal{V}_{1}$; the families $\mathcal{V}_{2}$ and $\mathcal{V}_{3}$ can be deduced from $\mathcal{V}_{1}$ by rotations of angles $\pm \frac{2 \pi}{3}$.

$$
\begin{align*}
& X(\theta)=A^{2}(-\left.\frac{1}{2}+\frac{\cos (2 \theta)}{6}\right)+A^{4}\left(\frac{19}{72}+\frac{107 \cos (2 \theta)}{432}+\frac{\cos (4 \theta)}{2160}\right) \\
&+A^{6}\left(-\frac{9241}{41472}+\frac{17761 \cos (2 \theta)}{103680}+\frac{349 \cos (4 \theta)}{388800}+\frac{\cos (6 \theta)}{115200}\right) \\
&+A^{8}\left(\frac{26178763}{74649600}+\frac{4912669 \cos (2 \theta)}{111974400}-\frac{60097 \cos (4 \theta)}{111974400}\right. \\
&\left.+\frac{409 \cos (6 \theta)}{22394880}+\frac{\cos (8 \theta)}{22394880}\right) \\
&+A^{10}\left(-\frac{11303303447}{26873856000}-\frac{36925058981 \cos (2 \theta)}{161243136000}-\frac{26246317 \cos (4 \theta)}{6718464000}\right. \\
&\left.+\frac{463349 \cos (6 \theta)}{35831808000}+\frac{137 \cos (8 \theta)}{1182449664000}\right)+\cdots \\
&+A^{20}\left(\frac{140898252923082652738578947}{27963744157874257920000000}+\cdots\right. \\
&\left.+\frac{14745542820149 \cos (20 \theta)}{13312230645358437943046307840000000}\right)+\mathrm{O}\left(A^{21}\right) \\
& Y(\theta)=A \cos ( \theta) \\
&+A^{3} \frac{\cos (3 \theta)}{48}+A^{5}\left(\frac{1657 \cos (3 \theta)}{34560}+\frac{17 \cos (5 \theta)}{103680}\right)  \tag{3.26}\\
&+A^{7}\left(\frac{120553 \cos (3 \theta)}{2764800}+\frac{53839 \cos (5 \theta)}{74649600}+\frac{71 \cos (7 \theta)}{74649600}\right) \\
&+A^{9}\left(\frac{24987797 \cos (3 \theta)}{3583180800}+\frac{3742409 \cos (5 \theta)}{2687385600}\right. \\
&\left.+\frac{64711 \cos (7 \theta)}{10749542400}+\frac{103 \cos (9 \theta)}{17915904000}\right)+\cdots \\
&+A^{21}\left(\frac{63416721855039974586904819141 \cos (3 \theta)}{157044387190621832478720000000}+\cdots\right. \\
&
\end{align*}
$$

$$
\begin{aligned}
& \left.+\frac{542127808389263 \cos (21 \theta)}{3904920098305141796626916966400000000}\right)+\mathrm{O}\left(A^{20}\right) \\
\omega^{2}=1-\frac{5}{6} A^{2} & +\frac{673}{864} A^{4}-\frac{18053}{69120} A^{6}+\frac{13759463}{17915904} A^{8}-\frac{16915570541}{16124313600} A^{10} \\
& +\frac{4774203682903}{5159780352000} A^{12}-\frac{87632953243730063}{33435376680960000} A^{14} \\
& +\frac{83706128400928836749}{24073471210291200000} A^{16}-\frac{61194528445910688474923}{12710792799033753600000} A^{18} \\
& +\frac{1772713703500289738506498277}{151004218452520992768000000} A^{20}+\mathrm{O}\left(A^{21}\right)
\end{aligned}
$$

We remark here that

$$
\begin{gather*}
X_{2 j}(\theta)=0  \tag{3.27}\\
Y_{2 j+1}(\theta)=0
\end{gather*} \quad \forall j \in \mathbb{N}
$$

(3) 'Circular' periodic family $\mathcal{C}$. This family contains in fact two circular families $\mathcal{C}^{+}$and $\mathcal{C}^{-}$having the same orbit but with two opposite senses of circulation.

$$
\begin{align*}
X(\theta)=A \cos (\theta) & -A^{2} \frac{\cos (2 \theta)}{3}+A^{4}\left(\frac{8 \cos (2 \theta)}{27}-\frac{\cos (4 \theta)}{135}\right)+A^{5} \frac{\cos (5 \theta)}{1620} \\
& +A^{6}\left(-\frac{214 \cos (2 \theta)}{405}+\frac{223 \cos (4 \theta)}{12150}\right)+A^{7}\left(-\frac{571 \cos (5 \theta)}{291600}+\frac{\cos (7 \theta)}{116640}\right) \\
& +A^{8}\left(\frac{64463 \cos (2 \theta)}{54675}-\frac{102373 \cos (4 \theta)}{2187000}-\frac{\cos (8 \theta)}{874800}\right) \\
& +A^{9}\left(\frac{592267 \cos (5 \theta)}{104976000}-\frac{341 \cos (7 \theta)}{8398080}\right)+\cdots \\
& +A^{20}\left(\frac{83816790509301709072526567 \cos (2 \theta)}{14400884480716800000000}\right. \\
& -\frac{178629639940427873315207767 \cos (4 \theta)}{6336389171515392000000000} \\
& -\frac{5086088610691877630528293 \cos (8 \theta)}{132895202223916154880000000} \\
& +\frac{3460611329192643548389 \cos (10 \theta)}{148468858734531329280000} \\
& +\frac{44998399011033141931 \cos (14 \theta)}{95885851629133813598208000000} \\
& +\frac{106900220735506097 \cos (16 \theta)}{85492629949052875760640000000} \\
& \left.+\frac{232217 \cos (20 \theta)}{151173284836324608000000}\right)+\mathrm{O}\left(A^{21}\right) \\
Y(\theta)= \pm[A & \sin (\theta)-A^{2} \frac{\sin (2 \theta)}{3}+A^{4}\left(\frac{-8 \sin (2 \theta)}{27}-\frac{\sin (4 \theta)}{135}\right)-A^{5} \frac{\sin (5 \theta)}{1620} \\
& +A^{6}\left(\frac{214 \sin (2 \theta)}{405}+\frac{223 \sin (4 \theta)}{12150}\right)+A^{7}\left(\frac{571 \sin (5 \theta)}{291600}+\frac{\sin (7 \theta)}{116640}\right) \\
& +A^{8}\left(\frac{-64463 \sin (2 \theta)}{54675}-\frac{102373 \sin (4 \theta)}{2187000}+\frac{\sin (8 \theta)}{874800}\right)
\end{align*}
$$

$$
\begin{aligned}
& +A^{9}\left(\frac{-592267 \sin (5 \theta)}{104976000}-\frac{341 \sin (7 \theta)}{8398080}\right)+\cdots \\
& +A^{20}\left(\frac{-83816790509301709072526567 \sin (2 \theta)}{14400884480716800000000}\right. \\
& -\frac{178629639940427873315207767 \sin (4 \theta)}{6336389171515392000000000} \\
& +\frac{5086088610691877630528293 \sin (8 \theta)}{1328952022239161548800000000} \\
& +\frac{3460611329192643548389 \sin (10 \theta)}{148468858734531329280000000} \\
& -\frac{44998399011033141931 \sin (14 \theta)}{95885851629133813598208000000} \\
& -\frac{106900220735506097 \sin (16 \theta)}{85492629949052875760640000000} \\
& \left.\left.+\frac{232217 \sin (20 \theta)}{151173284836324608000000}\right)+\mathrm{O}\left(A^{21}\right)\right] \\
& \omega^{2}=1+\frac{2}{3} A^{2}-\frac{16}{27} A^{4}+\frac{428}{405} A^{6}-\frac{257851}{109350} A^{8}+\frac{23207359}{3936600} A^{10}-\frac{12435443447}{787320000} A^{12} \\
& +\frac{113100657808997}{2550916800000} A^{14}-\frac{118199616016583611}{918330048000000} A^{16} \\
& +\frac{1393719688073359127771}{3636586990080000000} A^{18} \\
& -\frac{3352579515833898822938051}{2880176896143360000000} A^{20}+\mathrm{O}\left(A^{21}\right) .
\end{aligned}
$$

We notice in figures $1-3$ that numerical integration gives orbits which are indistinguishable to the naked eye from those obtained by the LP method. It is useful to show the figure grouping, for the energy $W=\frac{1}{8}$, the eight orbits computed by the LP method.

## 4. Study of the periods of the three main periodic families

### 4.1. Coefficients of the periods of the periodic familes

Since the energy is a constant of motion, we can determine the period of each periodic family in terms of $E$ in the following form

$$
\begin{equation*}
T=2 \pi \sum_{j=0}^{\infty} c_{j} E^{j} \tag{4.1}
\end{equation*}
$$

In tables $1-3$ we list, for $0 \leqslant j \leqslant 10$, the values $c_{j}$ of the three periodic families $\mathcal{R}, \mathcal{V}$ and $\mathcal{C}$.

### 4.2. Exact rectilinear families and their periods

Recall that we have found, by means of the LP method, one main 'rectilinear' family which in fact contains three 'rectilinear' familes, the 'horizontal' $\mathcal{H}(y=0)$ and the two 'oblique' $\mathcal{O}^{ \pm}(y= \pm \sqrt{3} x)$.

Table 1. Coefficients of the 'rectilinear' period.

| $j$ | $c_{j}$ |
| :---: | :---: |
| 0 | 1 |
| 1 | $\frac{5}{6}$ |
| 2 | $\frac{385}{144}$ |
| 3 | $\frac{85085}{7776}$ |
| 4 | $\frac{37182145}{746496}$ |
| 5 | $\frac{1078282205}{4478976}$ |
| 6 | 1169936192425 96758816 |
| 7 | $\frac{36220269467525}{5804752896}$ |
| 8 | 73201164593868025 |
| 9 | 190103424450275260925 |
| 10 | $\underline{24675424493645728868065}$ |

Table 2. Coefficients of the 'curvilinear' period.

| $j$ | $c_{j}$ |
| :---: | :---: |
| 0 | 1 |
| 1 | $\frac{5}{6}$ |
| 2 | $\frac{49}{144}$ |
| 3 | $-\frac{127967}{38880}$ |
| 4 | $-\frac{72840859}{3732480}$ |
| 5 | $-\frac{4799708427}{11974400}$ |
| 6 | 133960064084501 |
| 7 | $\frac{5197543150445477}{36279050500}$ |
| 8 |  |
| 9 | $\frac{78039878077739881278059}{37938650778419200000}$ |
| 10 | $-\frac{5315976060024828721532973}{4468638093410304000000}$ |

We will rigorously confirm this result by a direct calculation. Actually, a rectilinear solution $(x, y)$ can be searched in the form

$$
\begin{align*}
x(t) & =A \cdot \varphi(t)  \tag{4.2}\\
y(t) & =B \cdot \varphi(t)
\end{align*}
$$

Substituting (4.2) in the system (1.2) we obtain
$B=0 \quad$ or $\quad B=\sqrt{3} A \quad$ or $\quad B=-\sqrt{3} A \quad$ or $\quad \varphi=0$.
The case $\varphi=0$ corresponds to a trivial solution ( 0,0 ). The other cases represent the rectilinear families

$$
\begin{equation*}
y=0 \quad y=\sqrt{3} x \quad y=-\sqrt{3} x \tag{4.4}
\end{equation*}
$$

We shall now search the exact periods of these rectilinear families.
The Hamiltonian of the 'horizontal' family is given by

$$
\begin{equation*}
H=\frac{1}{2}\left(\dot{x}^{2}+x^{2}\right)-\frac{1}{3} x^{3} \tag{4.5}
\end{equation*}
$$

Table 3. Coefficients of the 'circular' period.

| $j$ | $c_{j}$ |
| :--- | :--- |
| 0 | 1 |
| 1 | $-\frac{1}{3}$ |
| 2 | $\frac{7}{9}$ |
| 3 | $-\frac{6559}{2430}$ |
| 4 | $\frac{324877}{29160}$ |
| 5 | $-\frac{110794621}{2187000}$ |
| 6 | $\frac{1157380170569}{4723920000}$ |
| 7 | $-\frac{87616855970353}{70858800000}$ |
| 8 | $\frac{437969924611388303}{68024448000000}$ |
| 9 | $-\frac{24971859775321935749203}{7273173980160000000}$ |
| 10 | $\frac{162836728825368274866249487}{872780877619200000000}$ |

The Hamiltonian of the 'oblique' families is given by

$$
\begin{equation*}
\frac{H}{4}=\frac{1}{2}\left(\dot{x}^{2}+x^{2}\right)+\frac{2}{3} x^{3} . \tag{4.6}
\end{equation*}
$$

The system is thus reduced to one degree of freedom and we can then apply a method based on the distributions $[8,9]$ to prove that the period of the three rectilinear families is rigorously equal to a Gauss hypergeometric series

$$
\begin{equation*}
\frac{T}{2 \pi}={ }_{2} F_{1}\left(\frac{1}{6}, \frac{5}{6} ; 1 ; 6 E\right)=\sum_{j=0}^{\infty} d_{j} E^{j} \tag{4.7}
\end{equation*}
$$

where the coefficients $d_{j}$ are given by

$$
\begin{equation*}
d_{j}=\frac{\left(\frac{1}{6}\right)_{j} \cdot\left(\frac{5}{6}\right)_{j} \cdot 6^{j}}{(j /!)^{2}} \tag{4.8}
\end{equation*}
$$

with

$$
\begin{align*}
& (a)_{0}=1 \\
& (a)_{j}=a(a+1) \ldots(a+j-1) \quad \text { if } j \geqslant 1 . \tag{4.9}
\end{align*}
$$

The comparison of the coefficients $c_{j}$ (table 1) with the above hypergeometric coefficients $d_{j}$ gives a perfect agreement. This therefore constitutes a check of the rectilinear periodic families obtained by the LP method.

### 4.3. Discussion of the convergence of the periods

We can easily prove that the radius of convergence of the hypergeometric series representing the rectilinear period is equal to 1 .

Concerning the circular period, we have applied the Burlirsh-Stoer algorithm [12, 13] to the ratios sequence $\left(\frac{c_{j}}{c_{j+1}}\right)$, where the coefficients $c_{j}$ are given in the table 3. This algorithm consists of accelerating the convergence of the ratios sequence by means of a new sequence $\left(b_{j}\right)$ such that $\lim _{j \rightarrow \infty} b_{j}=\lim _{j \rightarrow \infty} \frac{c_{j}}{c_{j+1}}=-\frac{1}{R}$. In table 4 we list the values $\frac{c_{j}}{c_{j+1}}$ and $b_{j}$ for $0 \leqslant j \leqslant 24$.

Table 4.

| $j$ | $\frac{c_{j}}{c_{j+1}}$ | $b_{j}$ |
| :---: | :---: | :---: |
| 0 | -3 | 0 |
| 1 | -0.428571428571 | -3 |
| 2 | -0.288 153681963 | -0.197418393587 |
| 3 | -0.242270 151472 | -0.110650287026 |
| 4 | -0.219918392969 | -0.153 161118297 |
| 5 | -0.206774219435 | -0.15 8798755784 |
| 6 | -0.198 143409350 | -0.157248276011 |
| 7 | -0.192 050131766 | -0.156 883596441 |
| 8 | -0.187522053866 | -0.156987228828 |
| 9 | -0.184026244856 | -0.156985919230 |
| 10 | -0.181 246586079 | -0.156987501421 |
| 11 | -0.178983861354 | -0.156976031160 |
| 12 | -0.177 106364807 | -0.156976 581339 |
| 13 | -0.175 523571818 | -0.156976 506560 |
| 14 | -0.174 171227949 | -0.156976428099 |
| 15 | $\overline{-0.173002473328 ~}$ | -0.156976477272 |
| 16 | -0.171982334723 | -0.156976484373 |
| 17 | -0.171084184281 | -0.156976485820 |
| 18 | -0.170287392204 | -0.156976486047 |
| 19 | -0.169 575729027 | -0.156976485 403 |
| 20 | -0.168936252471 | -0.156976486386 |
| 21 | -0.168358515761 | -0.156976486276 |
| 22 | -0.167833994111 | -0.156976486277 |
| 23 | -0.167355662362 | -0.156976486275 |
| 24 | -0.166917679256 | -0.156976486279 |

We can deduce that the sequence $\left(b_{j}\right)$ converges (numerically) to the value $-\frac{1}{R} \approx$ -0.156976486 , and consequently we obtain, with a good accuracy, the radius of convergence of the circular period

$$
\begin{equation*}
R \approx 6.37038083 \tag{4.10}
\end{equation*}
$$

On the other hand, the application of the Burlirsh-Stoer algorithm and other algorithms $\dagger$ (cf [13]) to the curvilinear period does not give any information about its radius of convergence. We still study this problem as well as the question of the nature of the series representing the circular and the curvilinear periods.

## 5. Comparison with the 'geometrical' method of Churchill-Pecelli-Rod [9]

The three authors of [9] have used a 'geometrical' method which consists of constructing the periodic orbits by exploiting the symmetries of the potential. Below we give the figure of the eight periodic orbits 'geometrically' constructed by these authors for an energy $E$ less then the escape energy $\frac{1}{6}$.

Comparison of the above figure with figure 4 leads to the following remarks.
(1) We obtain eight orbits by the two methods.
(2) Figures 4 and 5 present three rectilinear families $y=0, y= \pm \sqrt{3} x$.
(3) There is a good resemblance (forms and positions) between the 'curvilinear' orbits of figure 4 and those of figure 5 .

[^1]

Figure 4. Eight periodic orbits computed by the LP method. (Energy $E=\frac{1}{8}$.)


Figure 5. Eight periodic orbits 'geometrically' constructed by Churchill-Pecelli-Rod ( $\epsilon=$ $1, E<\frac{1}{6}$ ).
(4) Figures 4 and 5 present two 'circular' periodic orbits.

However, we notice a disagreement in the position of the circular orbits in figures 4 and 5 .
It is useful to mention that the three authors of [9] have conjectured the position of these circular orbits, not by using their 'geometrical' approach, but with 'numerical' explorations on which they have expressed some doubts. In our work, we have resolved this problem by giving not only the position of the eight periodic orbits at any time and for any sufficiently small energy $E$, but also their periods in terms of the energy $E$.

## 6. Conclusion

The main aim of this work is to show the interest of the LP method in the research of the periodic solutions of the Hamiltonian systems. Its importance lies in the fact that it can simultaneously be used as a means of enumeration of the main periodic families and also
as a tool for the determination of these as well as their periods in the form of powers series.
In this paper, we have successfully applied the LP method to the Hénon-Heiles nonintegrable Hamiltonian system. We have proven that this system admits three main periodic families in the neighbourhood of the origin: the 'rectilinear' $\mathcal{R}$, the curvilinear $\mathcal{V}$ and the 'circular' $\mathcal{C}$. We have also shown that the period of the 'rectilinear' family is rigorously equal to a Gauss hypergeometric series (see equation (4.7)).

By means of the computer algebra system 'Mathematica', we have computed with exact accuracy the periodic families as well as their periods to higher order $\mathrm{O}\left(A^{21}\right)$, where $A$ is a zeroth-order amplitude. We have also proven that the technique known as 'elimination of secular terms', on which the LP method is based, is mathematically equivalent to the 'alternative of Fredholm'. We have therefore tested the LP series using a numerical integration; the rectilinear families have also been checked by direct calculation. Finally we have compared our results with those of the 'geometrical' method of Churchill-Pecelli-Rod [9]. This comparison has led to a good agreement, but we have noticed a disagreement concerning the 'circular' family.

Let us mention that we have applied the LP method, in another work, to two nonintegrable Hamiltonian systems stemming from astronomy:

$$
\begin{align*}
& H=\frac{1}{2}\left(\dot{x}^{2}+\dot{y}^{2}+x^{2}+y^{2}\right) \epsilon x y^{2} \quad \text { the Barbanis-Contopoulos system. }  \tag{I}\\
& H=\frac{1}{2}\left(\dot{x}^{2}+\dot{y}^{2}+x^{2}+y^{2}\right)+\epsilon x^{2} y^{2} \quad \text { the Ollongren system. }
\end{align*}
$$

We have found that system (I) admits, in the neighbourhood of the origin, six main periodic families from which three are rectilinear $(y=0$ and $y= \pm \sqrt{2} x)$. We have also shown that the period of the 'horizontal' family $y=0$ is equal to $2 \pi$ and the period of the two oblique $y= \pm \sqrt{2} x$ is equal to a Gauss hypergeometric series $T=2 \pi_{2} F_{1}\left(\frac{1}{6}, \frac{5}{6} ; 1 ; 8 \epsilon^{2} E\right)$.

Concerning system (II), we have found that it admits, in the neighbourhood of the origin, six main periodic families from which four are rectilinear ( $x=0, y=0$ and $y= \pm x$ ). We have also shown that the period of the horizontal and vertical families $(y=0, x=0)$ is equal to $2 \pi$, whereas the period of the two oblique families is equal to a Gauss hypergeometric series $T=2 \pi F_{1}\left(\frac{1}{4}, \frac{3}{4} ; 1 ;-4 \epsilon E\right)$.

We finally emphasize that the LP method remains little exploited in other fields such as the theory of stability and bifurcation, and the research of quasiperiodic solutions of Hamiltonian systems.

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[^1]:    $\dagger$ Epsilon-algorithm, Padé approximants.

